

# Relativistic Viscous Fluids

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## Introduction

This talk will proceed in the following way, all of this work being done in Mathematica using xAct:

- We will present the stress-energy tensor for the perfect fluid.
- We will then show how we can get the fluid equations of motion for a perfect fluid using the principle of source conservation applied to the stress-energy tensor.
- We will present the stress-energy tensor for a viscous fluid.
- I will show the xAct implementation of this tensor.
- I will then show how we can get the fluid equations of motion for a viscous fluid using the principle of source conservation applied to the stress-energy tensor.

## The Stress-Energy Tensor for a Perfect Fluid

Given the pressure  $p$ , the density  $\rho$ , the metric tensor  $g_{\alpha\beta}$ , and the 4-velocity  $u^\alpha$ , we can write the stress energy tensor for a perfect fluid

$$T^{\alpha\beta} = p g_{\alpha\beta} + (\rho + p) u^\alpha u^\beta. \quad (1)$$

How do we tell Mathematica to write this? First we have to load xAct.

```
In[1]:= << xAct`xTensor` ;
<< xAct`xPert` ;
<< xAct`xCoba` ;
<< xAct`SymManipulator` ;
$DefInfoQ = False;
$CVVerbose = False;
```

---



## The Stress-Energy Tensor for a Perfect Fluid II

Then we define our spacetime manifold.

```
In[7]:= DefManifold[M, 4, {α, β, γ, λ, μ, ν}];
```

Then we define the abstract metric.

```
In[8]:= DefMetric[-1, g[-μ, -ν], CD, {";", "∇"}, PrintAs → "g"];
```

We now establish our chart. We will label it cb for coordinate basis

```
In[9]:= DefChart[cb, M, {0, 1, 2, 3}, {t[], r[], θ[], φ[]}]
```

Then we can specify how we want these written in output,

```
In[10]:= cb /: CIndexForm[0, cb] := "t";
```

```
In[11]:= cb /: CIndexForm[1, cb] := "r";
```

```
In[12]:= cb /: CIndexForm[2, cb] := "θ";
```

```
In[13]:= cb /: CIndexForm[3, cb] := "φ";
```



## The Stress-Energy Tensor for a Perfect Fluid III

We are using geometrized units in this work where  $G = 1$  and  $c = 1$ . We need to define the constant for the central mass,  $M$ .

```
In[14]:= DefConstantSymbol[M]
```

We will adopt the Schwarzschild metric.

```
In[16]:= MatrixForm[met = DiagonalMatrix[{1 - (2 M) / r[], (1 - (2 M) / r[])^-1, -r[]^2, 2 r[] Sin[θ[]]^2}]]
```

```
Out[16]//MatrixForm=
```

$$\begin{pmatrix} 1 - \frac{2M}{r} & 0 & 0 & 0 \\ 0 & \frac{1}{1 - \frac{2M}{r}} & 0 & 0 \\ 0 & 0 & -r^2 & 0 \\ 0 & 0 & 0 & 2 r \sin[\theta]^2 \end{pmatrix}$$





## The Stress-Energy Tensor for a Perfect Fluid IV

We then enter this into our coordinate basis.

```
In[17]:= MetricInBasis[g, -cb, met] // TableForm
```

```
Out[17]/TableForm=
```

$$\begin{array}{cccc}
 g_{tt} \rightarrow 1 - \frac{2M}{r} & g_{tr} \rightarrow 0 & g_{t\theta} \rightarrow 0 & g_{t\phi} \rightarrow 0 \\
 g_{rt} \rightarrow 0 & g_{rr} \rightarrow \frac{1}{1 - \frac{2M}{r}} & g_{r\theta} \rightarrow 0 & g_{r\phi} \rightarrow 0 \\
 g_{\theta t} \rightarrow 0 & g_{\theta r} \rightarrow 0 & g_{\theta\theta} \rightarrow -r^2 & g_{\theta\phi} \rightarrow 0 \\
 g_{\phi t} \rightarrow 0 & g_{\phi r} \rightarrow 0 & g_{\phi\theta} \rightarrow 0 & g_{\phi\phi} \rightarrow 2r \sin[\theta]^2
 \end{array}$$

We can show the symmetries.

```
In[18]:= TensorValues@g
```

```
Out[18]= FoldedRule[{{grt → gtr, gθt → gtθ, gθr → grθ, gφt → gtφ, gφr → grφ, gφθ → gθφ},
  {gtt → 1 - (2M)/r, gtr → 0, gtθ → 0, gtφ → 0, grr → 1 / (1 - (2M)/r),
  grθ → 0, grφ → 0, gθθ → -r2, gθφ → 0, gφφ → 2r Sin[θ]2}]
```



## The Stress-Energy Tensor for a Perfect Fluid V

We can also enter it directly using CTensor.

```
In[19]:= gm = CTensor[met, {-cb, -cb}];
```

```
In[20]:= SetCMetric[gm, -cb];
```

The  $tt$  component of the tensor can be displayed.

```
In[21]:= gm[{0, -cb}, {0, -cb}]
```

```
Out[21]= 1 - (2 M) / r
```

We then define the density field.

```
In[23]:= DefTensor[ρ[], M];
```

Here we define the 4-velocity.

```
In[246]:= u = CTensor[{1, 0, 0, 0}, {-cb}];
```



# The Stress-Energy Tensor for a Perfect Fluid VI

For the perfect fluid we need to define a pressure field.

```
In[25]:= DefTensor[p[], M];
```

The stress-energy tensor for the perfect fluid is,

```
In[247]:= Tf[α_, β_] := (ρ[] × u[-α] × u[-β]) - (p[] × u[-α] × u[-β]) + p[] × gm[-α, -β]
```

```
In[248]:= Tf[α, β]
```

Out[248]=

$-\rho + p \left(1 - \frac{2M}{r}\right) + \rho$	$0$	$0$	$0$
$0$	$\frac{p}{1 - \frac{2M}{r}}$	$0$	$0$
$0$	$0$	$-p r^2$	$0$
$0$	$0$	$0$	$2 p r \sin[\theta]^2$

$\alpha\beta$

We can specify a particular set of components.

```
In[249]:= Tf[{-θ, cb}, {-θ, cb}]
```

Out[249]=  $-\rho + p \left(1 - \frac{2M}{r}\right) + \rho$



## The Stress-Energy Tensor for a Perfect Fluid VII

We can make a command to do this,

```
In[250]:= SET[α_, β_] := Tf[{-α, cb}, {-β, cb}]
```

```
In[251]:= SET[0, 0]
```

```
Out[251]= -p + p (1 - 2M/r) + ρ
```

```
In[252]:= SET[0, 1]
```

```
Out[252]= 0
```

```
In[253]:= SETTable = Table[Row[{"T"Grid[{{α,β}}], "→", SET[α, β]}, {α, 0, 3}, {β, 0, 3}] // TableForm
```

```
Out[253]/TableForm=
```

$T_{\theta \ 0 \rightarrow} -p + p \left(1 - \frac{2M}{r}\right) + \rho$	$T_{\theta \ 1 \rightarrow} 0$	$T_{\theta \ 2 \rightarrow} 0$	$T_{\theta \ 3 \rightarrow} 0$
$T_{1 \ 0 \rightarrow} 0$	$T_{1 \ 1 \rightarrow} \frac{p}{1 - \frac{2M}{r}}$	$T_{1 \ 2 \rightarrow} 0$	$T_{1 \ 3 \rightarrow} 0$
$T_{2 \ 0 \rightarrow} 0$	$T_{2 \ 1 \rightarrow} 0$	$T_{2 \ 2 \rightarrow} -p r^2$	$T_{2 \ 3 \rightarrow} 0$
$T_{3 \ 0 \rightarrow} 0$	$T_{3 \ 1 \rightarrow} 0$	$T_{3 \ 2 \rightarrow} 0$	$T_{3 \ 3 \rightarrow} 2 p r \sin[\theta]^2$





## Using the Source Conservation Law to Derive the Equations of a Perfect Fluid

The source conservation law is often called the conservation of 4-momentum. Why the distinction? There is no general conservation of 4-momentum in general relativity, only in the source of the gravitation field.

We can write this conservation law

$$T^{\alpha\beta}_{;\gamma} = 0. \quad (2)$$

In our case, by (1) we can write this, in terms of the velocity of the fluid  $v$ ,

$$T^{00}_{,0} + T^{00}_{;j} = \partial_t \rho + \rho v^i_{;j} = 0. \quad (3)$$

Looking at this long enough, we will realize that it is the continuity equation.

We also have the spatial components

$$T^{ij}_{,0} + T^{ij}_{;j} = \partial_t \rho v^i + (\rho v^i v^j)_{;j} + p_{,i} = 0. \quad (4)$$

This is none other than the Euler equation for a fluid.



## Using the Source Conservation Law to Derive the Equations of a Perfect Fluid II

In general we can write the Christoffel symbols

```
In[32]:= Christoffel[CD, PDcb][α, -β, -γ]
```

```
Out[32]:= Γ[∇, D] α β γ
```

We can calculate the covariant derivative of the metric.

```
In[232]:= cdg = CovDofMetric[gm];
```

We can calculate the Christoffel symbols.

```
In[233]:= chr = Christoffel[cdg, PDcb];
```

We can determine a specific Christoffel symbol

```
In[234]:= chr[{0, cb}, {0, -cb}, {1, -cb}]
```

```
Out[234]:= -  $\frac{M}{2 M r - r^2}$ 
```

We can make a command to do this,

```
In[235]:= Γ[α_, β_, γ_] := chr[{α, cb}, {β, -cb}, {γ, -cb}]
```

```
In[237]:= Γ[0, 0, 1]
```

```
Out[237]:= -  $\frac{M}{2 M r - r^2}$ 
```

We can get all the components.

```
In[37]:= ChrisTable =
Table[Row[{"Γ"Grid[{{α}}]
Grid[{{n,β,γ}}"},"→",Γ[α,β,γ]},{α,0,3},{β,0,3},{γ,0,3}]] // TableForm
```

Out[37]/TableForm=

$\Gamma_{00}^0 \rightarrow 0$	$\Gamma_{10}^0 \rightarrow -\frac{M}{2Mr-r^2}$	$\Gamma_{20}^0 \rightarrow 0$	$\Gamma_{30}^0 \rightarrow 0$
$\Gamma_{00}^0 \rightarrow -\frac{M}{2Mr-r^2}$	$\Gamma_{11}^0 \rightarrow 0$	$\Gamma_{21}^0 \rightarrow 0$	$\Gamma_{31}^0 \rightarrow 0$
$\Gamma_{00}^0 \rightarrow 0$	$\Gamma_{12}^0 \rightarrow 0$	$\Gamma_{22}^0 \rightarrow 0$	$\Gamma_{32}^0 \rightarrow 0$
$\Gamma_{00}^0 \rightarrow 0$	$\Gamma_{13}^0 \rightarrow 0$	$\Gamma_{23}^0 \rightarrow 0$	$\Gamma_{33}^0 \rightarrow 0$
$\Gamma_{00}^1 \rightarrow \frac{M(2M-r)}{r^3}$	$\Gamma_{10}^1 \rightarrow 0$	$\Gamma_{20}^1 \rightarrow 0$	$\Gamma_{30}^1 \rightarrow 0$
$\Gamma_{00}^1 \rightarrow 0$	$\Gamma_{11}^1 \rightarrow \frac{M}{2Mr-r^2}$	$\Gamma_{21}^1 \rightarrow 0$	$\Gamma_{31}^1 \rightarrow 0$
$\Gamma_{00}^1 \rightarrow 0$	$\Gamma_{12}^1 \rightarrow 0$	$\Gamma_{22}^1 \rightarrow -2M+r$	$\Gamma_{32}^1 \rightarrow 0$
$\Gamma_{00}^1 \rightarrow 0$	$\Gamma_{13}^1 \rightarrow 0$	$\Gamma_{23}^1 \rightarrow 0$	$\Gamma_{33}^1 \rightarrow \frac{(2M-r)\text{Sin}[\theta]^2}{r}$
$\Gamma_{00}^2 \rightarrow 0$	$\Gamma_{10}^2 \rightarrow 0$	$\Gamma_{20}^2 \rightarrow 0$	$\Gamma_{30}^2 \rightarrow 0$
$\Gamma_{00}^2 \rightarrow 0$	$\Gamma_{11}^2 \rightarrow 0$	$\Gamma_{21}^2 \rightarrow \frac{1}{r}$	$\Gamma_{31}^2 \rightarrow 0$
$\Gamma_{00}^2 \rightarrow 0$	$\Gamma_{12}^2 \rightarrow \frac{1}{r}$	$\Gamma_{22}^2 \rightarrow 0$	$\Gamma_{32}^2 \rightarrow 0$
$\Gamma_{00}^2 \rightarrow 0$	$\Gamma_{13}^2 \rightarrow 0$	$\Gamma_{23}^2 \rightarrow 0$	$\Gamma_{33}^2 \rightarrow \frac{\text{Sin}[2\theta]}{r}$
$\Gamma_{00}^3 \rightarrow 0$	$\Gamma_{10}^3 \rightarrow 0$	$\Gamma_{20}^3 \rightarrow 0$	$\Gamma_{30}^3 \rightarrow 0$
$\Gamma_{00}^3 \rightarrow 0$	$\Gamma_{11}^3 \rightarrow 0$	$\Gamma_{21}^3 \rightarrow 0$	$\Gamma_{31}^3 \rightarrow \frac{1}{2r}$
$\Gamma_{00}^3 \rightarrow 0$	$\Gamma_{12}^3 \rightarrow 0$	$\Gamma_{22}^3 \rightarrow 0$	$\Gamma_{32}^3 \rightarrow \text{Cot}[\theta]$
$\Gamma_{00}^3 \rightarrow 0$	$\Gamma_{13}^3 \rightarrow \frac{1}{2r}$	$\Gamma_{23}^3 \rightarrow \text{Cot}[\theta]$	$\Gamma_{33}^3 \rightarrow 0$



## Using the Source Conservation Law to Derive the Equations of a Perfect Fluid III

We can write the covariant derivative of the stress-energy tensor.

```
In[38]:= cdstpf[ $\alpha$ _,  $\beta$ _,  $\gamma$ _] := cdg[- $\alpha$ ] [Tf[ $\beta$ ,  $\gamma$ ]]
```

We can write this,

```
In[265]:= cdstpf[ $\alpha$ ,  $\beta$ ,  $\gamma$ ] // ToCanonical // Simplify
```

Out[265]=

$$\begin{array}{cccc}
 \left( -\frac{2M}{r} (\mathcal{D}_t \rho) + \mathcal{D}_t \rho \right) & \frac{M(-\rho+\rho)}{(2M-r)r} & 0 & 0 \\
 -\frac{2M}{r} (\mathcal{D}_\theta \rho) + \mathcal{D}_\theta \rho & 0 & 0 & 0 \\
 -\frac{2M}{r} (\mathcal{D}_\phi \rho) + \mathcal{D}_\phi \rho & 0 & 0 & 0 \\
 \frac{M(-\rho+\rho)}{(2M-r)r} & \frac{\mathcal{D}_t p}{1-\frac{2M}{r}} & 0 & 0 \\
 0 & -\frac{r(\mathcal{D}_r p)}{2M-r} & 0 & 0 \\
 0 & \frac{\mathcal{D}_\theta p}{1-\frac{2M}{r}} & 0 & 0 \\
 0 & \frac{\mathcal{D}_\phi p}{1-\frac{2M}{r}} & 0 & 0 \\
 0 & 0 & -r^2 (\mathcal{D}_t p) & 0 \\
 0 & 0 & -r^2 (\mathcal{D}_r p) & 0 \\
 0 & 0 & -r^2 (\mathcal{D}_\theta p) & 0 \\
 0 & 0 & -r^2 (\mathcal{D}_\phi p) & 0 \\
 0 & 0 & 0 & 2r \sin^2[\theta] (\mathcal{D}_t p) \\
 0 & 0 & 0 & 2r \sin^2[\theta] (\mathcal{D}_r p) \\
 0 & 0 & 0 & 2r \sin^2[\theta] (\mathcal{D}_\theta p) \\
 0 & 0 & 0 & 2r \sin^2[\theta] (\mathcal{D}_\phi p)
 \end{array}$$

$\beta \gamma \alpha$

We can specify the specific components of this. Here we have the covariant derivative of the time-time component.

```
In[40]:= cdstpf[{0, cb}, {0, cb}, {0, cb}] // ToCanonical // Simplify
```

Out[40]=  $-\left( \frac{2M}{r} (\mathcal{D}_t \rho) \right) / r + \mathcal{D}_t \rho$

We can write a program for this,

```
In[41]:= cdsetpf[α_, β_, γ_] := cdstpf[{α, cb}, {β, cb}, {γ, cb}]
```

```
In[255]:= cdsetpf[0, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[255]} = -\frac{2 M (\mathcal{D}_t p)}{r} + \mathcal{D}_t \rho$$

```
In[256]:= cdsetpf[1, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[256]} = \begin{pmatrix} -\frac{2 M (\partial_x p)}{r} + \mathcal{D}_t \rho \\ -\frac{2 M (\partial_r p)}{r} + \mathcal{D}_r \rho \\ -\frac{2 M (\partial_\theta p)}{r} + \mathcal{D}_\theta \rho \\ -\frac{2 M (\partial_\phi p)}{r} + \mathcal{D}_\phi \rho \end{pmatrix} -1$$

```
In[257]:= cdsetpf[2, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[257]} = \begin{pmatrix} -\frac{2 M (\partial_x p)}{r} + \mathcal{D}_t \rho \\ -\frac{2 M (\partial_r p)}{r} + \mathcal{D}_r \rho \\ -\frac{2 M (\partial_\theta p)}{r} + \mathcal{D}_\theta \rho \\ -\frac{2 M (\partial_\phi p)}{r} + \mathcal{D}_\phi \rho \end{pmatrix} -2$$

```
In[258]:= cdsetpf[3, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[258]} = \begin{pmatrix} -\frac{2 M (\partial_x p)}{r} + \mathcal{D}_t \rho \\ -\frac{2 M (\partial_r p)}{r} + \mathcal{D}_r \rho \\ -\frac{2 M (\partial_\theta p)}{r} + \mathcal{D}_\theta \rho \\ -\frac{2 M (\partial_\phi p)}{r} + \mathcal{D}_\phi \rho \end{pmatrix} -3$$

## Using the Source Conservation Law to Derive the Equations of a Perfect Fluid IV

Putting together all of the derivatives of the time components

In[261]:= **tcdsetpft = Total[Table[cdsetpf[ $\alpha$ , 0, 0], { $\alpha$ , 0, 3}]] // ToCanonical // Simplify**

$$\text{Out[261]= } \frac{1}{r} \left( \begin{array}{c} -2 M (\mathcal{D}_t p) + r (\mathcal{D}_t \rho) \\ -2 M (\mathcal{D}_r p) + r (\mathcal{D}_r \rho) \\ -2 M (\mathcal{D}_\theta p) + r (\mathcal{D}_\theta \rho) \\ -2 M (\mathcal{D}_\phi p) + r (\mathcal{D}_\phi \rho) \end{array} \right) -3 + \begin{array}{c} -2 M (\mathcal{D}_t p) + r (\mathcal{D}_t \rho) \\ -2 M (\mathcal{D}_r p) + r (\mathcal{D}_r \rho) \\ -2 M (\mathcal{D}_\theta p) + r (\mathcal{D}_\theta \rho) \\ -2 M (\mathcal{D}_\phi p) + r (\mathcal{D}_\phi \rho) \end{array} -2 + \begin{array}{c} -2 M (\mathcal{D}_t p) + r (\mathcal{D}_t \rho) \\ -2 M (\mathcal{D}_r p) + r (\mathcal{D}_r \rho) \\ -2 M (\mathcal{D}_\theta p) + r (\mathcal{D}_\theta \rho) \\ -2 M (\mathcal{D}_\phi p) + r (\mathcal{D}_\phi \rho) \end{array} -1 - 2 M (\mathcal{D}_t p) + r (\mathcal{D}_t \rho)$$

Setting this to 0 will give us the continuity equation for a perfect fluid in Schwarzschild spacetime. This is how (3) is adapted to our specific geometry.



# Using the Source Conservation Law to Derive the Equations of a Perfect Fluid V

In order to get the equations of momentum we need to apply a projection tensor.

We begin by defining the projection tensor.

```
In[433]:= P[α_, β_] := gm[-α, -β] + u[-α] × u[-β]
```

```
In[434]:= P[α, β]
```

Out[434]=

$2 - \frac{2M}{r}$	0	0	0
0	$\frac{1}{1 - \frac{2M}{r}}$	0	0
0	0	$-r^2$	0
0	0	0	$2r \sin[\theta]^2$

$\alpha\beta$

We can specify a particular set of components.

```
In[142]:= P[{-θ, cb}, {-θ, cb}] // Simplify
```

Out[142]=  $2 - \frac{2M}{r}$

We can make a command to do this,

```
In[143]:= PT[α_, β_] := P[{-α, cb}, {-β, cb}] // Simplify
```

```
In[144]:= PT[0, 0]
```

```
Out[144]= 2 -  $\frac{2M}{r}$ 
```

```
Out[145]/TableForm=
```

$P_{0 \ 0} \rightarrow 2 - \frac{2M}{r}$	$P_{0 \ 1} \rightarrow 0$	$P_{0 \ 2} \rightarrow 0$	$P_{0 \ 3} \rightarrow 0$
$P_{1 \ 0} \rightarrow 0$	$P_{1 \ 1} \rightarrow \frac{1}{1 - \frac{2M}{r}}$	$P_{1 \ 2} \rightarrow 0$	$P_{1 \ 3} \rightarrow 0$
$P_{2 \ 0} \rightarrow 0$	$P_{2 \ 1} \rightarrow 0$	$P_{2 \ 2} \rightarrow -r^2$	$P_{2 \ 3} \rightarrow 0$
$P_{3 \ 0} \rightarrow 0$	$P_{3 \ 1} \rightarrow 0$	$P_{3 \ 2} \rightarrow 0$	$P_{3 \ 3} \rightarrow 2r \sin[\theta]^2$

## Using the Source Conservation Law to Derive the Equations of a Perfect Fluid VI

So we then write,

```
In[481]:= mom[ $\alpha$ _,  $\beta$ _,  $\gamma$ _,  $\mu$ _,  $\nu$ _] := P[ $\mu$ ,  $\nu$ ]  $\times$  cdstpf[ $\alpha$ ,  $\beta$ ,  $\gamma$ ] // ToCanonical // Simplify
```

```
In[482]:= mom[ $\alpha$ ,  $\beta$ ,  $\gamma$ ,  $\mu$ ,  $\nu$ ]
```



## The Stress-Energy Tensor for a Viscous Fluid I

It is important to realize that for a Non-Perfect Fluid the stress-energy tensor will have the form of the stress-energy tensor plus some additional terms,

$$T^{\alpha\beta} = \text{Stress - Energy Tensor : Perfect Fluid} + \text{Viscous Contributions} + \text{The Generation of Energy Fluxes.} \quad (5)$$

We will adopt the Eckert frame for our presentation as this is most generally applicable (though not for some many-body collision cases).

The stress-energy tensor for the Eckert frame of a viscous fluid exhibiting heat flow, having an anisotropic stress tensor  $\pi^{\alpha\beta}$ , the viscous bulk pressure  $\Pi$ , a projection tensor  $P^{\alpha\beta}$ , the heat-flux 4-vector  $q^\alpha$ ,

$$T^{\alpha\beta} = \rho u^\alpha u^\beta + \pi^{\alpha\beta} + (p + \Pi) P^{\alpha\beta} + q^\alpha u^\beta + q^\beta u^\alpha. \quad (6)$$

$$P^{\alpha\beta} = g^{\alpha\beta} + u^\alpha u^\beta. \quad (7)$$

$$\Pi = p - p_{\text{eq}}. \quad (8)$$

We also have the heat flux 4-vector,

$$q^\alpha = (0, q^{\hat{i}}) \quad (9)$$

where the hat represents spatial components in the comoving frame of the fluid.

# The Stress-Energy Tensor for a Viscous Fluid II

We then establish our equilibrium pressure.

```
In[435]:= DefConstantSymbol [peq]
```

Then we define the viscous bulk pressure.

```
In[436]:= Π[] := p[] - peq
```

The heat flux 4 - vector is.

```
In[438]:= q = CTensor[{0, q1, q2, q3}, {-cb}];
```

```
In[441]:= c = CTensor[{
  Sqrt[1/(1 - 2M/r)], 0, 0, 0}, {-cb}];
```

```
In[442]:= DefTensor[S[], M];
```

```
In[444]:= σ[α_, β_] := Sqrt[3] S[] (
  c[-α] × c[-β] - 1/3 (u[-α] × u[-β] - gm[-α, -β]))
```

```
In[445]:= σ[α, β]
```

Out[445]=

$\sqrt{3} \left( \frac{1}{1 - \frac{2M}{r}} - \frac{2M}{3r} \right) S$	0	0	0
0	$\frac{S}{\sqrt{3} \left( 1 - \frac{2M}{r} \right)}$	0	0
0	0	$-\frac{r^2 S}{\sqrt{3}}$	0
0	0	0	$\frac{2 r S \sin[\theta]^2}{\sqrt{3}}$

$\alpha\beta$

## The Stress-Energy Tensor for a Viscous Fluid III

```
In[453]:= Tvf[ $\alpha_*$ ,  $\beta_*$ ] :=
   $\rho$  [ $\alpha$ ]  $\times$  u[- $\alpha$ ]  $\times$  u[- $\beta$ ] +  $\sigma$ [ $\alpha$ ,  $\beta$ ] + ( $\rho$  [ $\alpha$ ] +  $\Pi$  [ $\alpha$ ]) P[ $\alpha$ ,  $\beta$ ] + q[- $\alpha$ ]  $\times$  u[- $\beta$ ] + q[- $\beta$ ]  $\times$  u[- $\alpha$ ]
```

```
In[455]:= Tvf[ $\alpha$ ,  $\beta$ ] // ToCanonical
```

```
Out[455]= 
$$\begin{pmatrix} 0 & q_1 & q_2 & q_3 \\ q_1 & 0 & 0 & 0 \\ q_2 & 0 & \rho e q & r^2 - 2 p & r^2 - \frac{r^2 s}{\sqrt{3}} & 0 \\ q_3 & 0 & 0 & 0 & 0 & 0 \end{pmatrix} \alpha \beta$$

```

We can specify a particular set of components.

```
In[456]:= Tvf[{- $\theta$ , cb}, {- $\theta$ , cb}]
```

```
Out[456]= 
$$(-\rho e q + 2 p) \left( 2 - \frac{2 M}{r} \right) + \sqrt{3} \left( \frac{1}{1 - \frac{2 M}{r}} - \frac{2 M}{3 r} \right) S + \rho$$

```

We can make a command to do this,

```
In[457]:= SETV[ $\alpha_*$ ,  $\beta_*$ ] := Tvf[{- $\alpha$ , cb}, {- $\beta$ , cb}]
```

```
In[459]:= SETV[ $\theta$ ,  $\theta$ ]
```

```
Out[459]= 
$$(-\rho e q + 2 p) \left( 2 - \frac{2 M}{r} \right) + \sqrt{3} \left( \frac{1}{1 - \frac{2 M}{r}} - \frac{2 M}{3 r} \right) S + \rho$$

```



## The Stress-Energy Tensor for a Viscous Fluid IV

```
In[460]:= SETVTable = Table[Row[{"T"Grid[{{α,β}}, "→", SETV[α, β]]], {α, 0, 3}, {β, 0, 3}] // TableForm
```

```
Out[460]//TableForm=
```

$$\begin{array}{lll}
 T_{0 \ 0} \rightarrow (-\text{peq} + 2 \text{p}) \left(2 - \frac{2M}{r}\right) + \sqrt{3} \left(\frac{1}{1 - \frac{2M}{r}} - \frac{2M}{3r}\right) S + \rho & T_{0 \ 1} \rightarrow q1 & T_{0 \ 2} \rightarrow q2 \\
 T_{1 \ 0} \rightarrow q1 & T_{1 \ 1} \rightarrow \frac{-\text{peq} + 2 \text{p}}{1 - \frac{2M}{r}} + \frac{S}{\sqrt{3} \left(1 - \frac{2M}{r}\right)} & T_{1 \ 2} \rightarrow 0 \\
 T_{2 \ 0} \rightarrow q2 & T_{2 \ 1} \rightarrow 0 & T_{2 \ 2} \rightarrow -(-\text{peq} + 2 \text{p}) r^2 \\
 T_{3 \ 0} \rightarrow q3 & T_{3 \ 1} \rightarrow 0 & T_{3 \ 2} \rightarrow 0
 \end{array}$$

## Using the Source Conservation Law to Derive the Equations of a Viscous Fluid I



We can write the covariant derivative of the stress-energy tensor.

```
In[468]:= cdstpfv[α_, β_, γ_] := cdg[-α][Tvf[β, γ]]
```

We can write this,

```
In[469]:= cdstpfv[α, β, γ] // ToCanonical // Simplify
```

Out[469]=

$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} \mathcal{D}_r q_1 \\ -\frac{q_2}{r} + \mathcal{D}_\theta q_1 \\ -\frac{q_3}{2r} + \mathcal{D}_\phi q_1 \end{pmatrix}$	$\begin{pmatrix} \frac{\mathcal{D}_t q_2}{(2M-r)r} + \mathcal{D}_r q_2 \\ 2M q_1 - q_1 r + \mathcal{D}_\theta q_2 \\ -q_3 \cot[\theta] + \mathcal{D}_\phi q_2 \end{pmatrix}$	$\begin{pmatrix} \frac{\mathcal{D}_t q_3}{4M-2r} + \mathcal{D}_r q_3 \\ -q_3 \cot[\theta] + \mathcal{D}_\phi q_3 \\ 0 \end{pmatrix}$
$\begin{pmatrix} \mathcal{D}_r q_1 \\ -\frac{q_2}{r} + \mathcal{D}_\theta q_1 \\ -\frac{q_3}{2r} + \mathcal{D}_\phi q_1 \end{pmatrix}$	$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} \frac{M q_2}{2M r - r^2} \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} \frac{M q_3}{2M r - r^2} \\ 0 \\ 0 \\ 0 \end{pmatrix}$
$\begin{pmatrix} \mathcal{D}_t q_2 \\ \frac{q_2}{(2M-r)r} + \mathcal{D}_r q_2 \\ 2M q_1 - q_1 r + \mathcal{D}_\theta q_2 \\ -q_3 \cot[\theta] + \mathcal{D}_\phi q_2 \end{pmatrix}$	$\begin{pmatrix} \frac{M q_2}{2M r - r^2} \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$
$\begin{pmatrix} \mathcal{D}_t q_3 \\ \frac{q_3}{4M-2r} + \mathcal{D}_r q_3 \\ -q_3 \cot[\theta] + \mathcal{D}_\phi q_3 \end{pmatrix}$	$\begin{pmatrix} \frac{M q_3}{2M r - r^2} \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$	$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$

$\beta \gamma \alpha$

We can specify the specific components of this. Here we have the covariant derivative of the time-time component.

```
In[470]:= cdstpfv[{0, cb}, {0, cb}, {0, cb}] // ToCanonical // Simplify
```

Out[470]=

$$\left(4 - \frac{4M}{r}\right) (\mathcal{D}_t p) + \frac{\left(-\frac{2M}{r} - \frac{3r}{2M-r}\right) (\mathcal{D}_t S)}{\sqrt{3}} + \mathcal{D}_t \rho$$

We can write a program for this,

```
In[471]:= cdsetpfv[α_, β_, γ_] := cdstpfv[{α, cb}, {β, cb}, {γ, cb}]
```

```
In[473]:= cdsetpfv[0, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[473]} = \left(4 - \frac{4M}{r}\right) (\mathcal{D}_t p) + \frac{\left(-\frac{2M}{r} - \frac{3r}{2M-r}\right) (\mathcal{D}_t S)}{\sqrt{3}} + \mathcal{D}_t \rho$$

```
In[474]:= cdsetpfv[1, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[474]} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix} - 1$$

```
In[475]:= cdsetpfv[2, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[475]} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix} - 2$$

```
In[476]:= cdsetpfv[3, 0, 0] // ToCanonical // Simplify
```

$$\text{Out[476]} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix} - 3$$

## Using the Source Conservation Law to Derive the Equations of a Viscous Fluid II

Putting together all of the derivatives of the time components

In[478]:= `tcdsetpftv = Total[Table[cdsetpfv[α, 0, 0], {α, 0, 3}]] // ToCanonical // Simplify`

Out[478]= 
$$\begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}_{-3} + \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}_{-2} + \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}_{-1} + 4 (\mathcal{D}_t \rho) - \frac{4 M (\mathcal{D}_t p)}{r} + \frac{\sqrt{3} (\mathcal{D}_t S)}{1 - \frac{2M}{r}} - \frac{2 M (\mathcal{D}_t S)}{\sqrt{3} r} + \mathcal{D}_t \rho$$

Setting this to 0 will give us the continuity equation for a viscous fluid in Schwarzschild spacetime.

## Using the Source Conservation Law to Derive the Equations of a Viscous Fluid III

In order to get the equations of momentum we need to apply a projection tensor.

```
In[484]:= momv[ $\alpha$ _,  $\beta$ _,  $\gamma$ _,  $\mu$ _,  $\nu$ _] := P[ $\mu$ ,  $\nu$ ]  $\times$  cdstpfv[ $\alpha$ ,  $\beta$ ,  $\gamma$ ] // ToCanonical // Simplify
```

```
In[485]:= mom[ $\alpha$ ,  $\beta$ ,  $\gamma$ ,  $\mu$ ,  $\nu$ ]
```





Once again, you can read the components off the lists to get the specific components. More work on this needs to be done to extract features like the vorticity tensor.



## References

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**Thank You!**